

Improvement of the Determination of the WIMP Mass from Direct Dark Matter Detection Data

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in collaboration with M. Drees

based on JCAP 0806, 012 (arXiv:0803.4477 [hep-ph])

Determining the WIMP mass model-independently

Correcting the systematic deviation of the reconstructed WIMP mass

Summary

Determining the WIMP mass model-independently

- Differential event rate for elastic WIMP-nucleus scattering

$$\frac{dR}{dQ} = \mathcal{A} F^2(Q) \int_{v_{\min}}^{v_{\text{esc}}} \left[\frac{f_1(v)}{v} \right] dv$$

Here

$$v_{\min} = \alpha \sqrt{Q}$$

is the minimal incoming velocity of incident WIMPs that can deposit the recoil energy Q in the detector.

$$\mathcal{A} \equiv \frac{\rho_0 \sigma_0}{2m_\chi m_{r,N}^2} \quad \alpha \equiv \sqrt{\frac{m_N}{2m_{r,N}^2}} \quad m_{r,N} = \frac{m_\chi m_N}{m_\chi + m_N}$$

ρ_0 : WIMP density near the Earth

σ_0 : total cross section ignoring the form factor suppression

$F(Q)$: elastic nuclear form factor

$f_1(v)$: one-dimensional velocity distribution of halo WIMPs

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Astrophysics

Particle Physics

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Determining the WIMP mass model-independently

- Determining the moments of the WIMP velocity distribution

$$\langle v^n \rangle = \alpha^n \left[\frac{2Q_{\text{thre}}^{1/2} r_{\text{thre}}}{F^2(Q_{\text{thre}})} + I_0 \right]^{-1} \left[\frac{2Q_{\text{thre}}^{(n+1)/2} r_{\text{thre}}}{F^2(Q_{\text{thre}})} + (n+1)I_n \right]$$

$$I_n = \sum_a \frac{Q_a^{(n-1)/2}}{F^2(Q_a)}$$

$$r_{\text{thre}} = \left(\frac{dR}{dQ} \right)_{Q=Q_{\text{thre}}}$$

[M. Drees and CLS, JCAP 0706, 011]

Determining the WIMP mass model-independently

- Determining the moments of the WIMP velocity distribution

$$\langle v^n \rangle = \alpha^n \left[\frac{2Q_{\text{thre}}^{1/2} r_{\text{thre}}}{F^2(Q_{\text{thre}})} + l_0 \right]^{-1} \left[\frac{2Q_{\text{thre}}^{(n+1)/2} r_{\text{thre}}}{F^2(Q_{\text{thre}})} + (n+1)l_n \right]$$

$$l_n = \sum_a \frac{Q_a^{(n-1)/2}}{F^2(Q_a)}$$

$$r_{\text{thre}} = \left(\frac{dR}{dQ} \right)_{Q=Q_{\text{thre}}}$$

[M. Drees and CLS, JCAP 0706, 011]

- Determining the WIMP mass

$$m_X = \frac{\sqrt{m_X m_Y} - m_X \mathcal{R}_n}{\mathcal{R}_n - \sqrt{m_X / m_Y}}$$

$$\mathcal{R}_n \equiv \frac{\alpha_Y}{\alpha_X}$$

$$= \left[\frac{2Q_{\text{thre},X}^{(n+1)/2} r_{\text{thre},X} / F_X^2(Q_{\text{thre},X}) + (n+1)l_{n,X}}{2Q_{\text{thre},X}^{1/2} r_{\text{thre},X} / F_X^2(Q_{\text{thre},X}) + l_{0,X}} \right]^{1/n} \left(X \longrightarrow Y \right)^{-1} \quad (n \neq 0)$$

[CLS and M. Drees, arXiv:0710.4296]

Determining the WIMP mass model-independently

- Spin-independent (SI) WIMP-nucleus cross section (**neutralino**)

$$\sigma_0^{\text{SI}} \simeq \left(\frac{4}{\pi}\right) m_{r,N}^2 A^2 |f_p|^2 \quad f_p: \text{effective WIMP - proton coupling}$$

- Determining the **WIMP mass**

$$m_X^{\text{SI}} = \frac{(m_X/m_Y)^{5/2} m_Y - m_X \mathcal{R}_\sigma}{\mathcal{R}_\sigma - (m_X/m_Y)^{5/2}}$$

$$\mathcal{R}_\sigma = \frac{\mathcal{E}_Y}{\mathcal{E}_X} \left[\frac{2Q_{\text{thre},X}^{1/2} r_{\text{thre},X} / F_X^2(Q_{\text{thre},X}) + I_{0,X}}{2Q_{\text{thre},Y}^{1/2} r_{\text{thre},Y} / F_Y^2(Q_{\text{thre},Y}) + I_{0,Y}} \right]$$

- Neglecting Q_{thre}

$$\mathcal{R}_\sigma = \frac{\mathcal{E}_Y}{\mathcal{E}_X} \left(\frac{I_{0,X}}{I_{0,Y}} \right)$$

$$\sigma \left(m_X^{\text{SI}} \right) = \frac{\mathcal{R}_\sigma (m_X/m_Y)^{5/2} |m_X - m_Y|}{\left[\mathcal{R}_\sigma - (m_X/m_Y)^{5/2} \right]^2} \left[\frac{\sigma^2(I_{0,X})}{I_{0,X}^2} + \frac{\sigma^2(I_{0,Y})}{I_{0,Y}^2} \right]^{1/2}$$

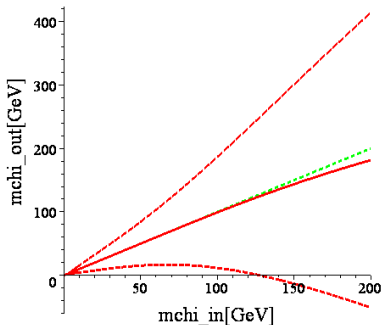
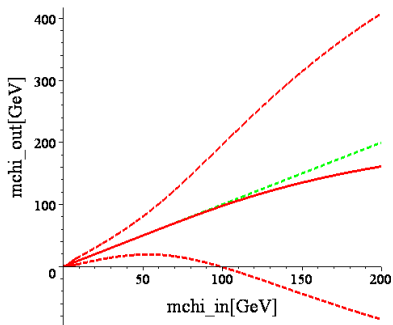
[M. Drees and CLS, JCAP 0806, 012]

Determining the WIMP mass model-independently

- Reconstructed WIMP mass $m_\chi(n=1)$ and m_χ^{SI}
(1 – 200 keV, $^{76}\text{Ge} + ^{28}\text{Si}$, 25 + 25 / 50 + 50 events)

$Q_{\text{max}} = 200$ keV, $Q_{\text{min}} = 1$ keV, $n = 1$, 25 + 25 events, Ge-76 + Si-28

$Q_{\text{max}} = 200$ keV, $Q_{\text{min}} = 1$ keV, 50 + 50 events, Ge-76 + Si-28



[CLS and M. Drees, arXiv:0710.4296]

- A smaller deviation, but a larger statistical error!

Correcting the systematic deviation of the reconstructed WIMP mass

- Matching Q_{\max}

$$Q_{\max, \text{Si}} = \left(\frac{\alpha_{\text{Ge}}}{\alpha_{\text{Si}}} \right)^2 Q_{\max, \text{Ge}} \quad \Leftarrow \quad v_{\max} = \alpha \sqrt{Q_{\max}}$$

- Algorithmic χ^2 fit

$$\chi^2 = \sum_{i,j} (f_{i,X} - f_{i,Y}) C_{ij}^{-1} (f_{j,X} - f_{j,Y})$$

where

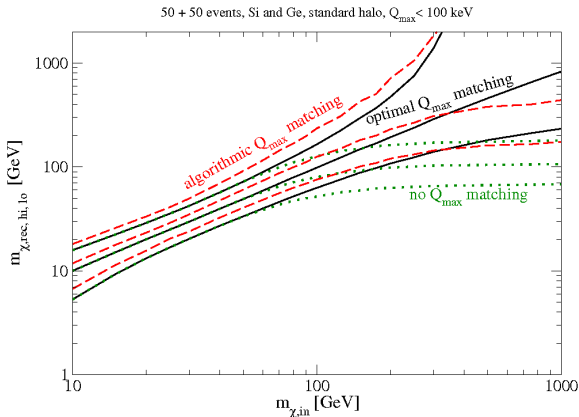
$$f_{i,X} = \alpha_X^i \left[\frac{2Q_{\min,X}^{(i+1)/2} r_X(Q_{\min,X}) / F_X^2(Q_{\min,X}) + (i+1)l_{i,X}}{2Q_{\min,X}^{1/2} r_X(Q_{\min,X}) / F_X^2(Q_{\min,X}) + l_{0,X}} \right] \left(\frac{1}{300 \text{ km/s}} \right)^i$$

$$f_{n_{\max}+1,X} = \mathcal{E}_X \left[\frac{A_X^2}{2Q_{\min,X}^{1/2} r_X(Q_{\min,X}) / F_X^2(Q_{\min,X}) + l_{0,X}} \right] \left(\frac{\sqrt{m_X}}{m_X + m_X} \right)$$

$$C_{ij} = \text{cov}(f_{i,X}, f_{j,X}) + \text{cov}(f_{i,Y}, f_{j,Y})$$

Correcting the systematic deviation of the reconstructed WIMP mass

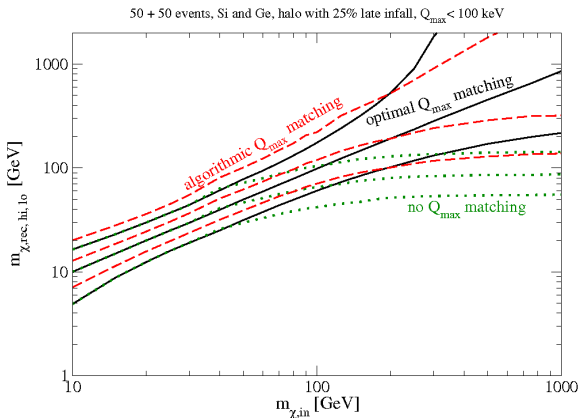
□ Reconstructed WIMP mass

($Q_{\max} < 100$ keV, $^{76}\text{Ge} + ^{28}\text{Si}$, 50 + 50 events)

[M. Drees and CLS, JCAP 0806, 012]

Correcting the systematic deviation of the reconstructed WIMP mass

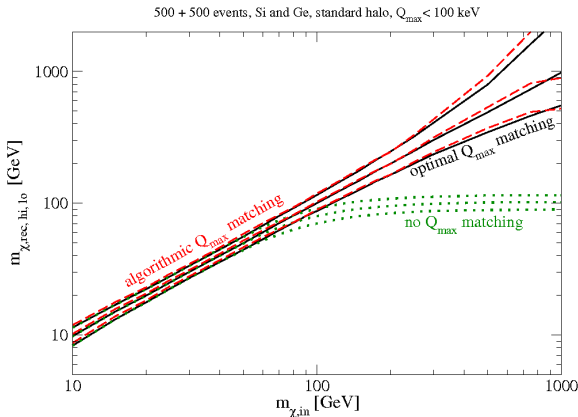
□ Reconstructed WIMP mass

($Q_{\max} < 100$ keV, $^{76}\text{Ge} + ^{28}\text{Si}$, 50 + 50 events, late infall compo.)

[M. Drees and CLS, JCAP 0806, 012]

Correcting the systematic deviation of the reconstructed WIMP mass

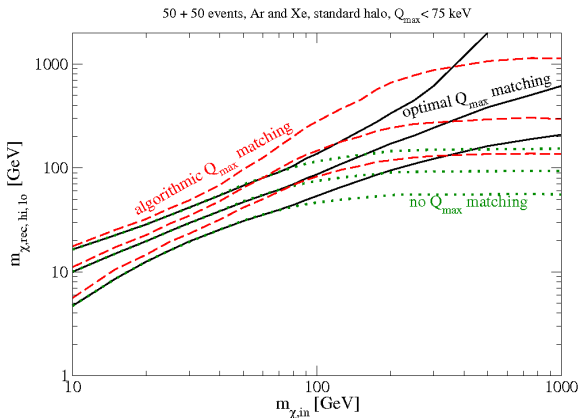
□ Reconstructed WIMP mass

($Q_{\max} < 100$ keV, $^{76}\text{Ge} + ^{28}\text{Si}$, 500 + 500 events)

[M. Drees and CLS, JCAP 0806, 012]

Correcting the systematic deviation of the reconstructed WIMP mass

□ Reconstructed WIMP mass

($Q_{\max} < 75$ keV, $^{40}\text{Ar} + ^{136}\text{Xe}$, 50 + 50 events)

[M. Drees and CLS, JCAP 0806, 012]

Summary

- Once two or more experiments **with different detector materials** obtain positive WIMP signals, we can determine the WIMP mass.
- Our methods are **model-independent** and require **only measured recoil energies**.
- The deviation of the reconstructed WIMP mass from the true one **can be corrected by matching more suitable cut-off energies**.
- With **~ 100 keV** maximal measuring energies and **2×50 events**, a WIMP mass of 50 GeV can be estimated with **an error of $\sim 35\%$** .